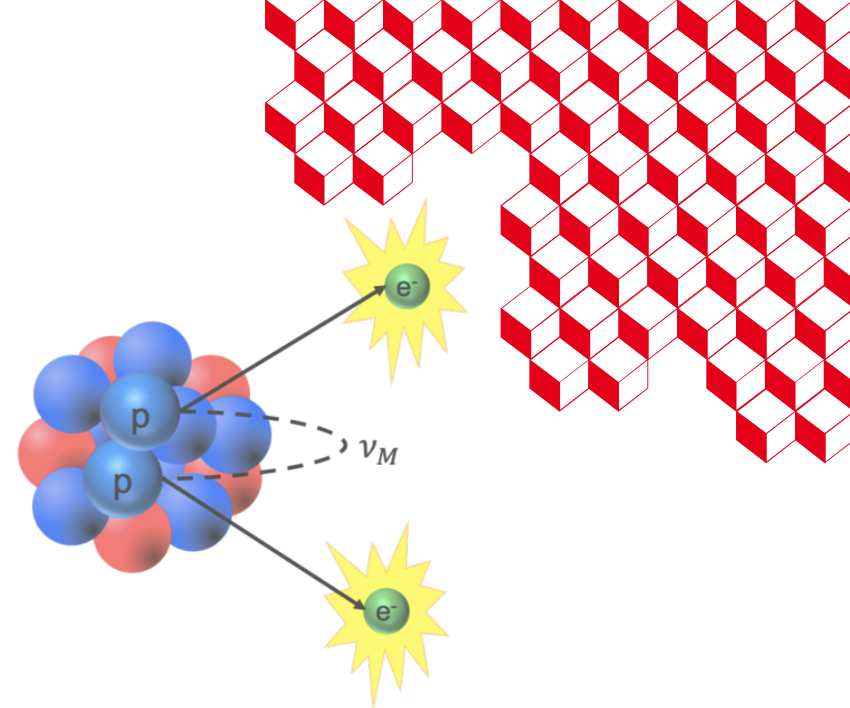




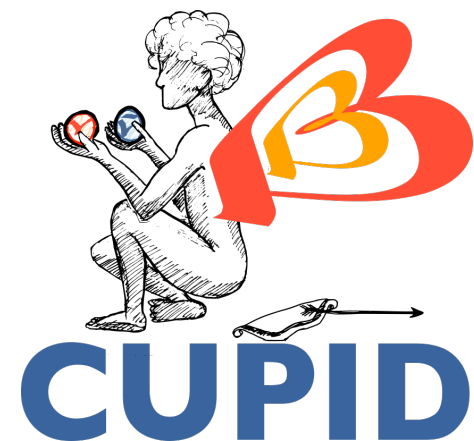
CUPID: the next generation $0\nu\beta\beta$ bolometric experiment

Claudia Nones

On behalf of the CUPID collaboration

XVIII
International Conference on
Topics in Astroparticle and
Underground Physics 2023
28.08. - 01.09.2023
University of Vienna

SCROLL



Outline

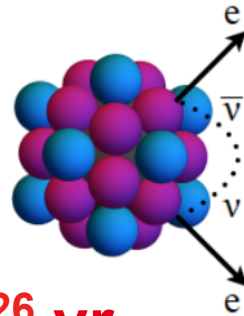
- Double beta decay → promising isotopes
- The bolometric technique → from CUORE to CUPID
- CUPID precursors and CUPID structure
- CUPID experimental tests
- Background and sensitivity
- Conclusions

Nuclear Double Beta Decay

- Double Beta Decay is the **rarest nuclear weak process**
- It takes place between **two even-even isobars**

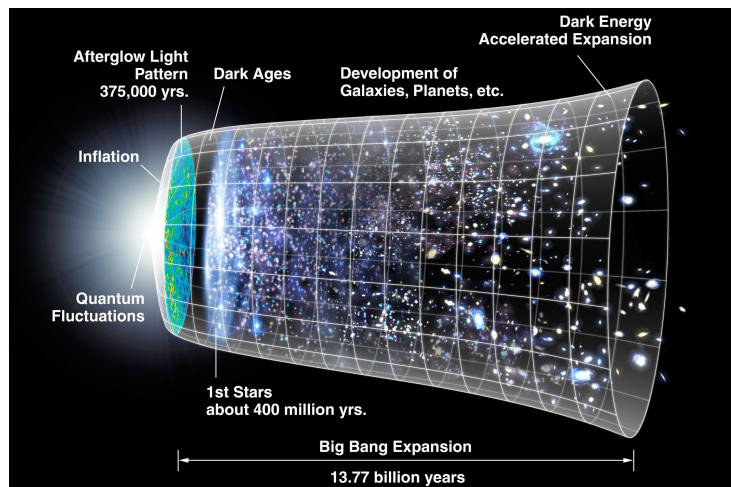
$$(A, Z) \rightarrow (A, Z+2) + 2e^- + (...)$$

If only electrons and nothing else:



Half-life larger $> 10^{25}$ yr - 10^{26} yr

- Need to find single events in a ton of isotope x year(s) of exposure!
- 3×10^{-11} Bq/kg vs 15 Bq/banana !
- We go to extreme length to limit ubiquitous radioactivity



$\times 10^{15}$

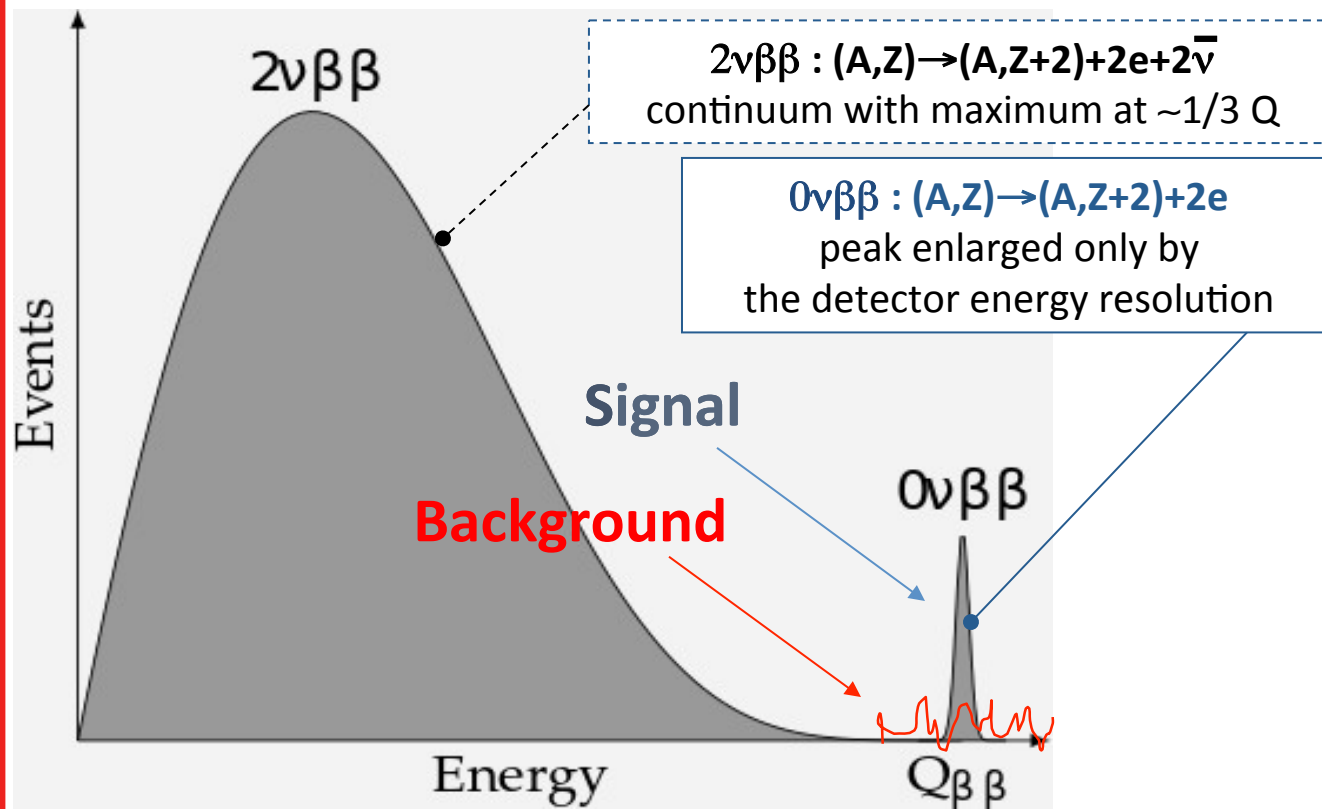
Very hard, but the reward is great:

- Lepton number non conservation
- Creation of matter with no antimatter counterpart
- Majorana nature of neutrinos
- Neutrino mass scale

Nuclear Double Beta Decay

Sum energy spectrum of the two electrons

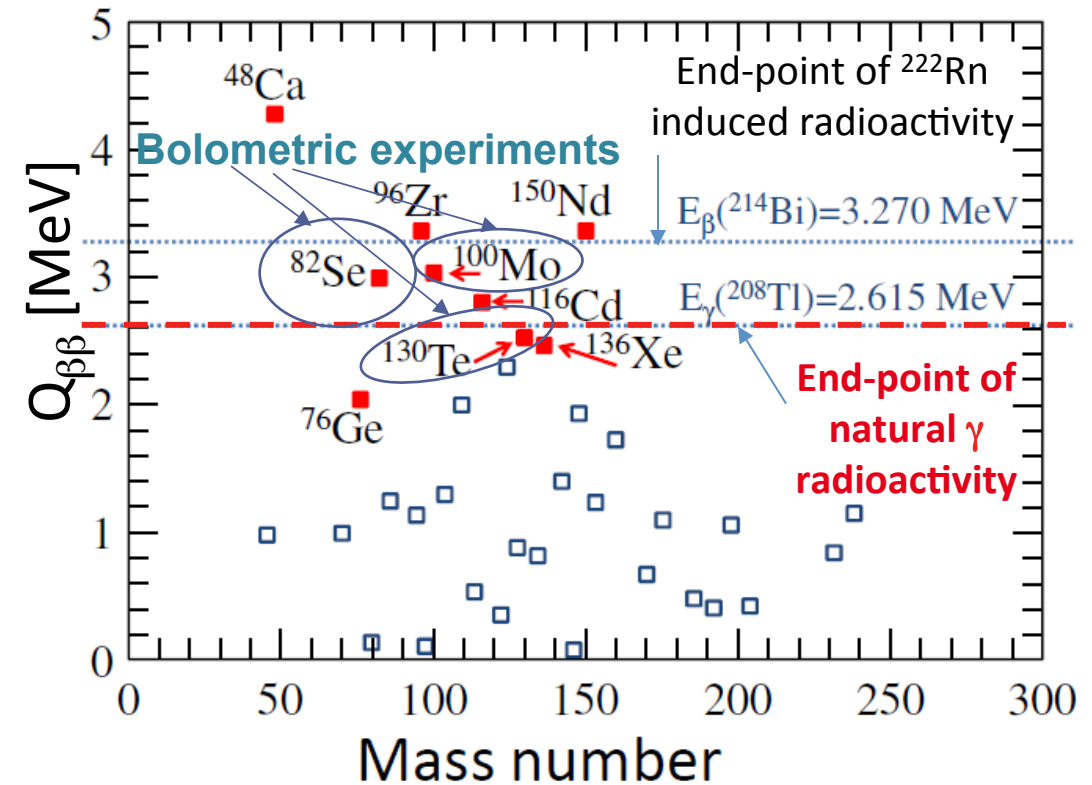
$Q_{\beta\beta}$: energy available for the products



Possible for 35 nuclei

Only 9 are **experimentally relevant**

$Q_{\beta\beta} \sim 2\text{-}3 \text{ MeV}$ for the most promising candidates



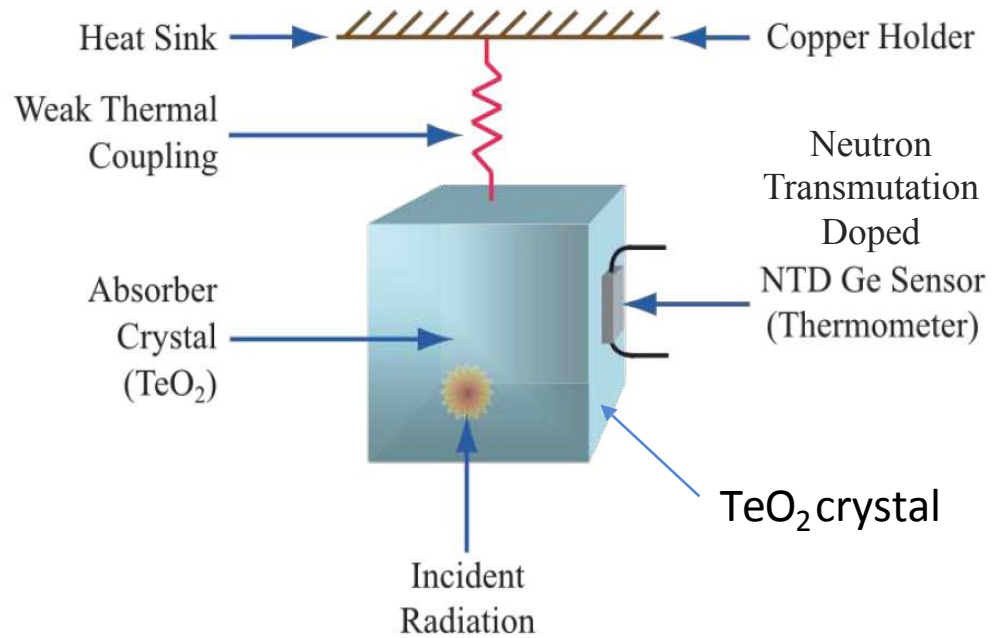
The bolometric technique: CUORE



CUORE ^{130}Te

pure thermal detector
(bolometer)

$T \sim 10 \text{ mK}$



No PID

$Q = 2527 \text{ keV} < 2615 \text{ keV}$

No enrichment: ^{130}Te I.A. $\sim 34\%$

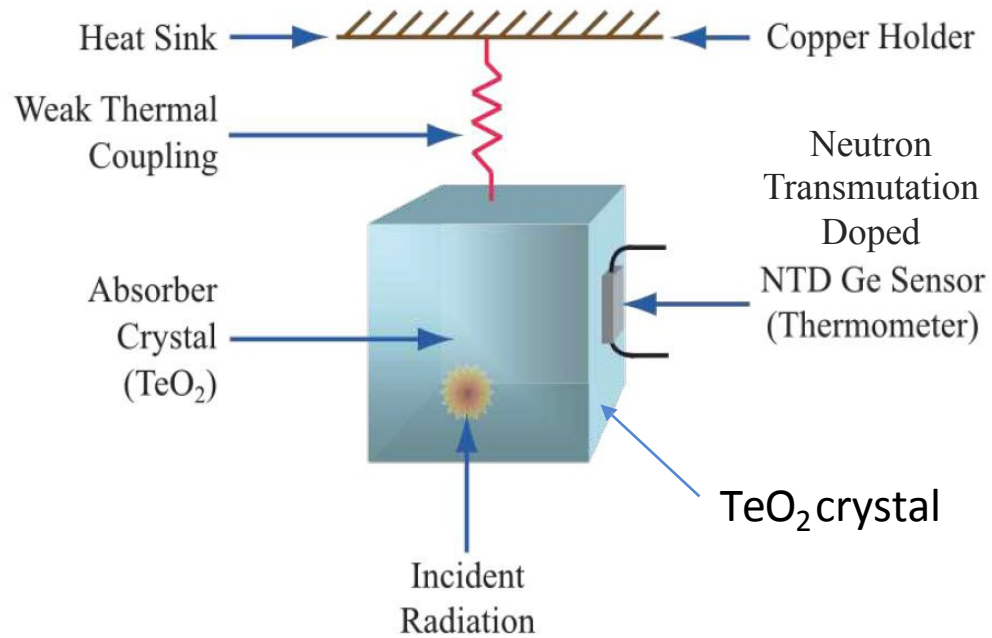
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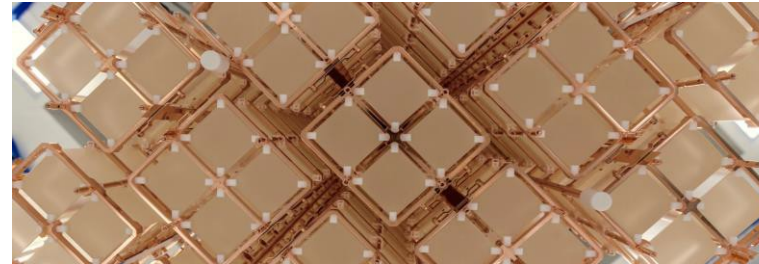
$Q = 2527 \text{ keV} < 2615 \text{ keV}$

No enrichment: ^{130}Te I.A. $\sim 34\%$

CUORE - Cryogenic Underground Observatory for Rare Events

- Located in Gran Sasso, Italy
- Main objective: $0\nu\beta\beta$ in ^{130}Te
- 988 TeO_2 crystals, $5 \times 5 \times 5 \text{ cm}^3$ each
- Total mass: 742 kg TeO_2 (natural Te)
- ^{130}Te mass: 206 kg
- Current analysed exposure: 2023 kg y
- $T_{1/2}^{0\nu} > 3.33 \times 10^{25} \text{ yr}$ at 90% C.I.
- $m_{\beta\beta} < 75 - 260 \text{ meV}$ at 90% C.I.

New released results @TAUP2023



α background $\rightarrow b \sim 10^{-2} \text{ c}/(\text{keV kg y})$

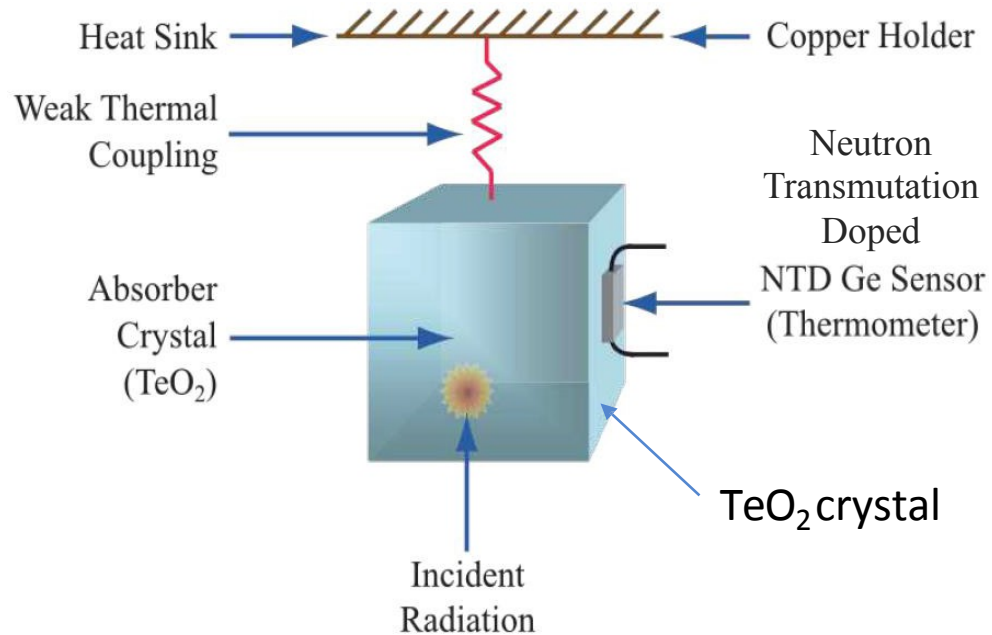
γ background $\rightarrow b \sim 10^{-3} \text{ c}/(\text{keV kg y})$

The bolometric technique: from CUORE to CUPID

CUORE ^{130}Te

pure thermal detector
(bolometer)

$T \sim 10 \text{ mK}$

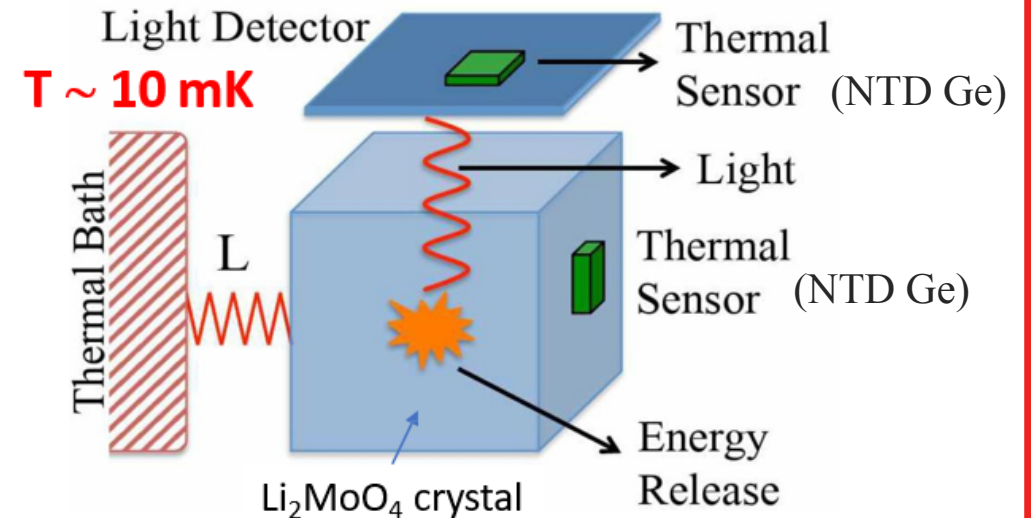


No PID

$Q = 2527 \text{ keV} < 2615 \text{ keV}$

No enrichment: ^{130}Te I.A. $\sim 34\%$

CUPID ^{100}Mo
heat + light
(scintillating bolometer)



~~α background~~ $\leftarrow$ **PID**

~~γ background~~ $\leftarrow$ **$Q = 3034 \text{ keV}$**
 $> 2615 \text{ keV}$

Goal: $b \sim 10^{-4} \text{ c}/(\text{keV kg y})$

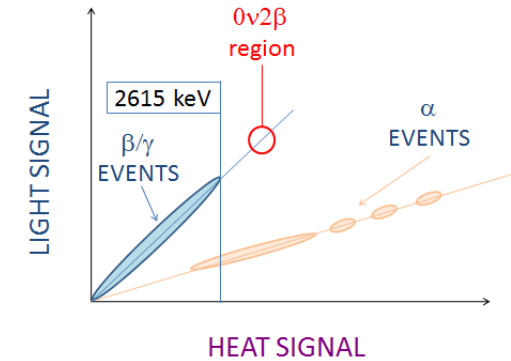
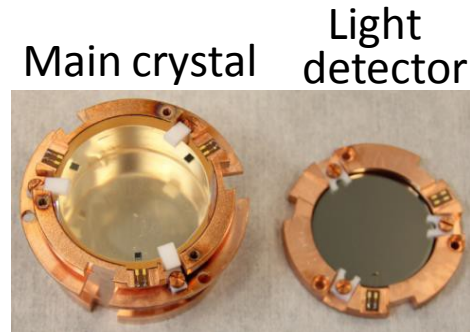
Enrichment in ^{100}Mo is possible on large scale

CUPID precursors

Scintillating bolometers



PID: α particle rejection



CUPID-0 – Zn^{82}Se $Q_{2\beta} = 2998 \text{ keV}$

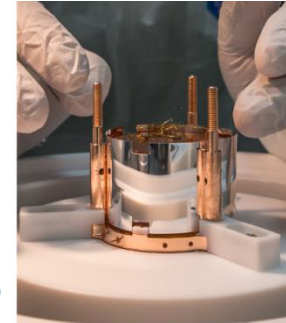
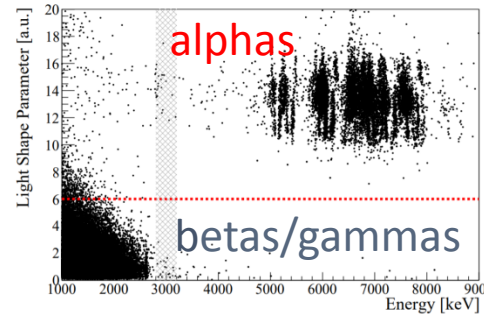
(evolution of **LUCIFER** )

First running demonstrator

24 crystals (enriched in ^{82}Se) – 5.28 kg ^{82}Se

Best limit on ^{82}Se : $T_{1/2} > 4.7 \times 10^{24} \text{ y}$

Energy resolution: **$\sim 23 \text{ keV FWHM}$**



LNGS – Italy $b = 3.5 \times 10^{-3} \text{ counts/(keV}\cdot\text{kg}\cdot\text{yr)}$
Useful information for the CUPID background model

Direct proof that α 's dominate background above 2.6 MeV

CUPID-Mo – $\text{Li}_2^{100}\text{MoO}_4$ $Q_{2\beta} = 3034 \text{ keV}$

(evolution of **LUMINEU** )

Physics data taking: April 2019 – June 2020

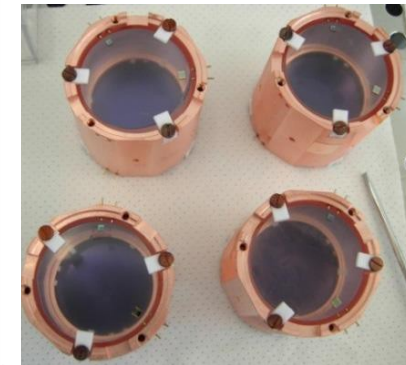
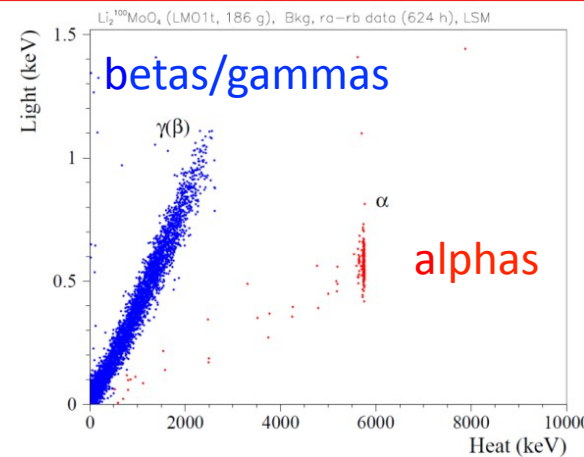
20 crystals (enriched in ^{100}Mo) – 2.34 kg ^{100}Mo

Energy resolution: **$\sim 5\text{--}7 \text{ keV FWHM}$**

Best limit on ^{100}Mo : $T_{1/2} > 1.8 \times 10^{24} \text{ y}$

Full α rejection

Radiopure crystals: $\text{U/Th} \leq 1 \mu\text{Bq/kg}$



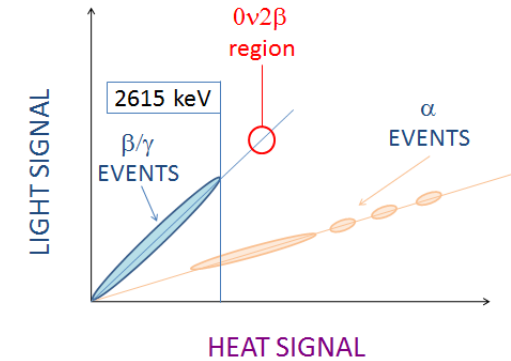
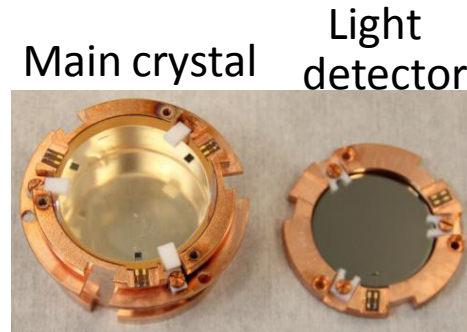
LSM – France
 $b = 2.7 \times 10^{-3} \text{ counts/(keV}\cdot\text{kg}\cdot\text{yr)}$
In non-optimized cryostat

CUPID precursors

Scintillating bolometers



PID: α particle rejection



CUPID-0 – Zn^{82}Se $Q_{2\beta} = 2998 \text{ keV}$

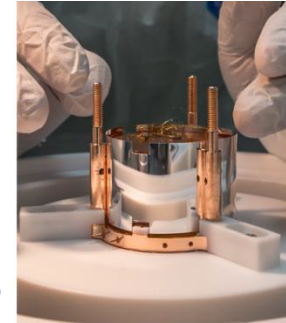
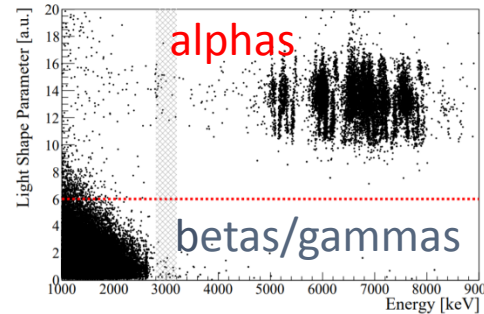
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24 crystals (enriched in ^{82}Se) – 5.28 kg ^{82}Se

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Energy resolution: **$\sim 23 \text{ keV FWHM}$**



LNGS – Italy $b = 3.5 \times 10^{-3}$
counts/(keV·kg·yr)

Useful information for the CUPID background model

Direct proof that α 's dominate background above 2.6 MeV

CUPID-Mo – $\text{Li}_2^{100}\text{MoO}_4$ $Q_{2\beta} = 3034 \text{ keV}$

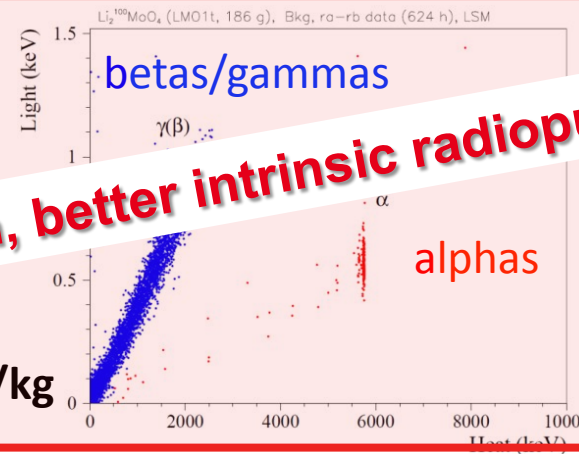
(evolution of **LUMINEU** )

Physics data taking: April 2019 – June 2020

20 crystals (enriched in ^{100}Mo) – 2.34 kg ^{100}Mo

Energy resolution: **$\sim 5-7 \text{ keV FWHM}$**

Best limit on ^{100}Mo : $T_{1/2} > 1.0 \times 10^{26} \text{ y}$



LSM – France

$b = 2.7 \times 10^{-3}$

In non-optimized cryostat

Choice for CUPID: better energy resolution, better intrinsic radiopurity, mature crystal technology

Radiopure crystals: U/Th $\leq 1 \mu\text{Bq/kg}$

The CUPID collaboration



7 countries
~180 members



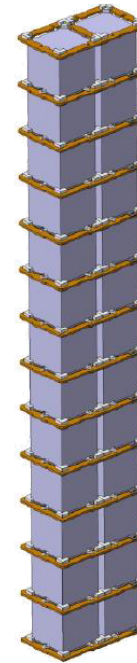
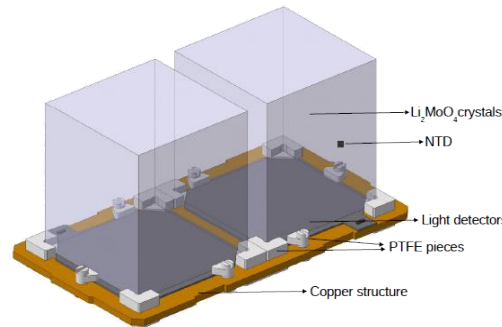
The CUPID experiment in a nutshell

CUPID pre-CDR *arXiv:1907.09376* upgrade to CDR ongoing

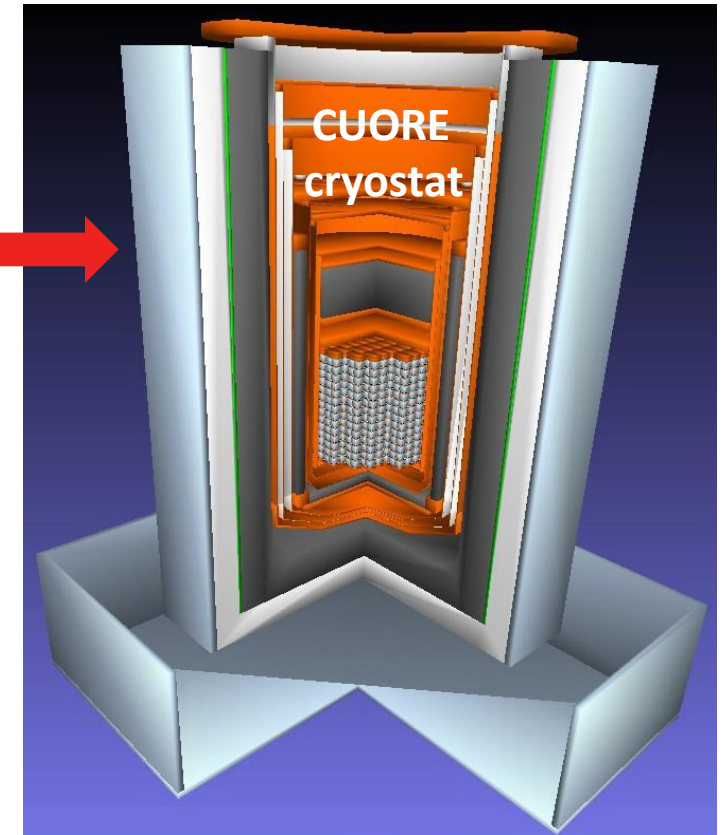
- Single module:
 - 2 x $\text{Li}_2^{100}\text{MoO}_4$ **45×45×45 mm** – **~280 g**
 - 2 x Ge light detectors
- 57 towers of 14 floors with 2 crystals each - **1596 crystals**
- ~240 kg of ^{100}Mo** with >95% enrichment ($\sim 1.6 \times 10^{27}$ ^{100}Mo atoms)
- 1710 Ge light detector** (CUPID-Mo, CUPID-0 basic technology + Neganov-Trofimov-Luke effect)

Baseline design

Gravity stacked structure
(innovative approach with respect
to CUORE and CUPID precursors)



Ton-scale array of high-resolution cryogenic calorimeters to search for $0\nu\beta\beta$ and other rare events



CUPID objectives

- Energy resolution at $Q_{\beta\beta}$: ≤ 5 keV FWHM
- Light signal: 0.3 keV/MeV
- Light detector baseline resolution: ≤ 100 eV RMS (for PID)
- Light detector timing resolution: ≤ 0.17 ms (for pile-up)
- Background index: $\leq 10^{-4}$ counts/(kg keV yr)
- $0\nu\beta\beta$ half-life exclusion sensitivity (90% C.L.): 1.4×10^{27} yr

Test of the tower structure

A prototype single CUPID tower was cooled down in July and October 2022 in LNGS

The tower was installed in the Cuoricino-CUORE0-CUPID-0 cryostat

- **Primary goal:** validate the tower structure in terms of assembly procedure, **detector temperature values and distribution**
- **Secondary goals:** Analysis of **detector performance** - Study of the **crystal radiopurity** – Dependence of detector behavior on **glue type, crystal origin, light-detector coating** method

14-floor tower run

- 28 LMO crystals
- 30 Ge light detectors

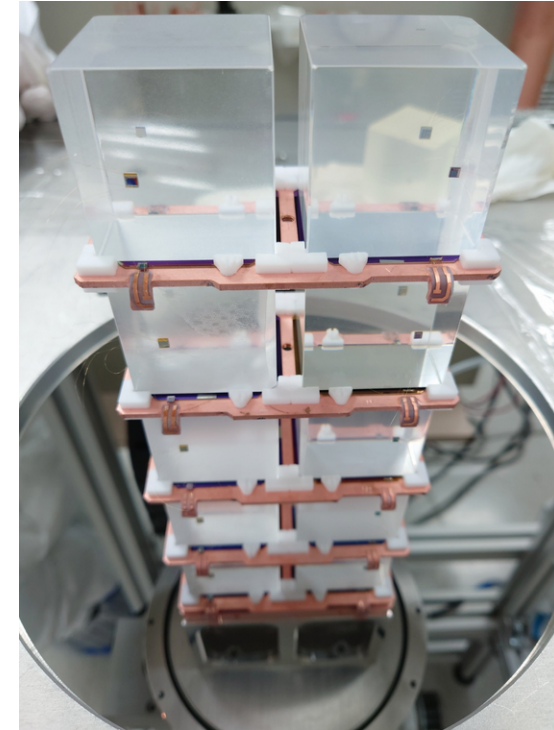
The primary goal is achieved - All the channels cool down without problems with a reasonably narrow temperature distribution and no dependence on the position in the tower

→ **The thermal scheme of the tower is validated**

LMO: 7-8 keV FWHM @ $Q_{\beta\beta}$

LD: 180 eV median RMS

New test in the coming months

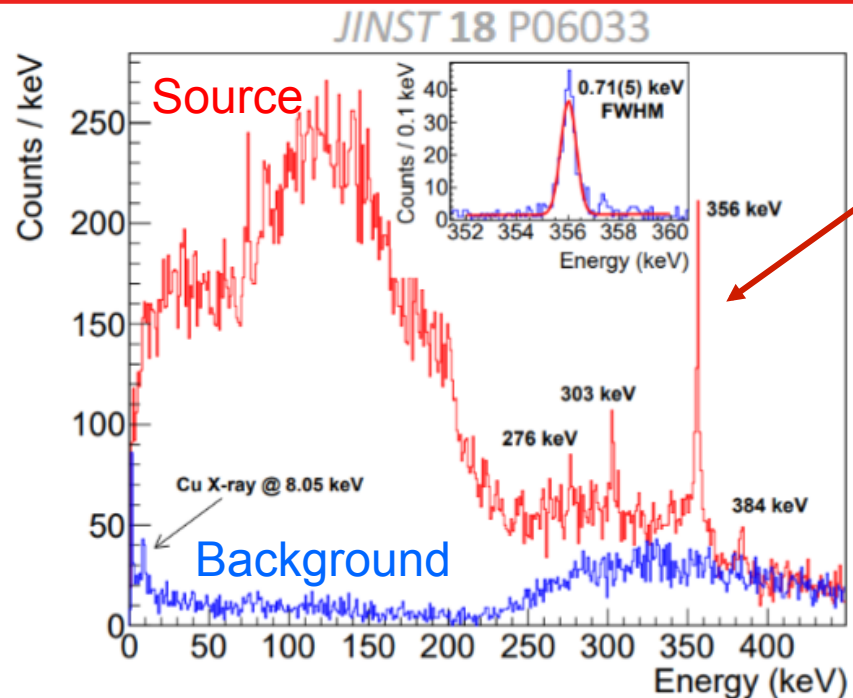
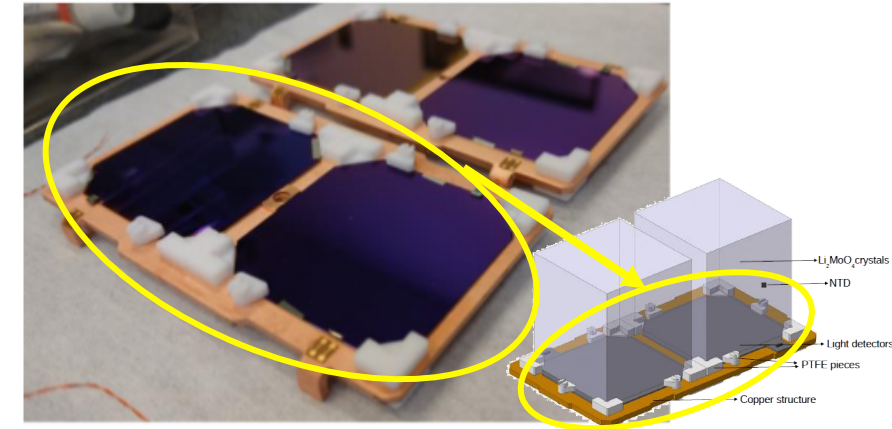


Test of the light-detector structure

The **validation of the new CUPID-baseline light-detector assembly** performed in a pulse-tube cryostat at IJCLab (Orsay)

→ Similar vibrational noise as in CUORE/CUPID cryostat

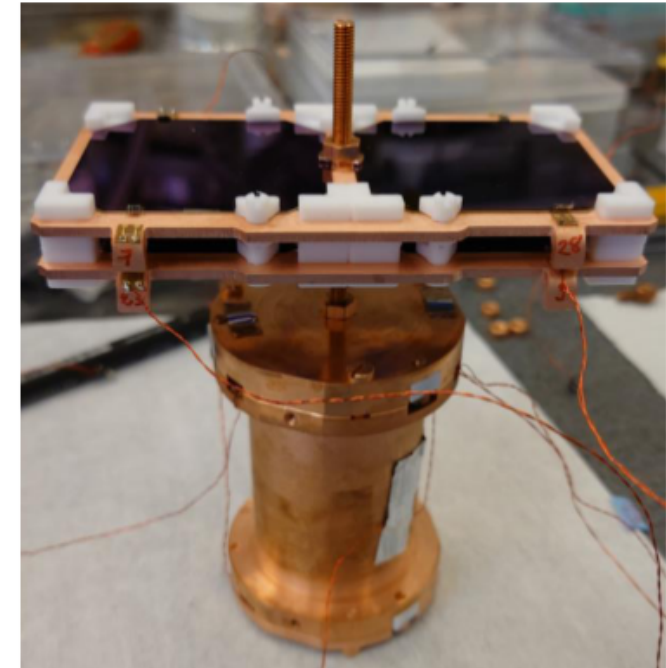
- 2 Cu frames with 2 light detectors each (0.5 mm in thickness) mounted on top of each other (CUPID configuration)
- Check of the bolometric performance
- **Baseline energy resolution 70-90 eV RMS** (CUPID goal for PID met)



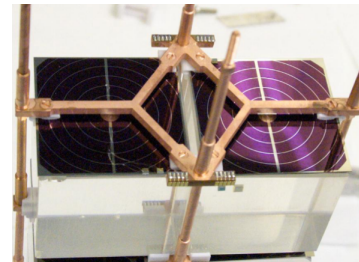
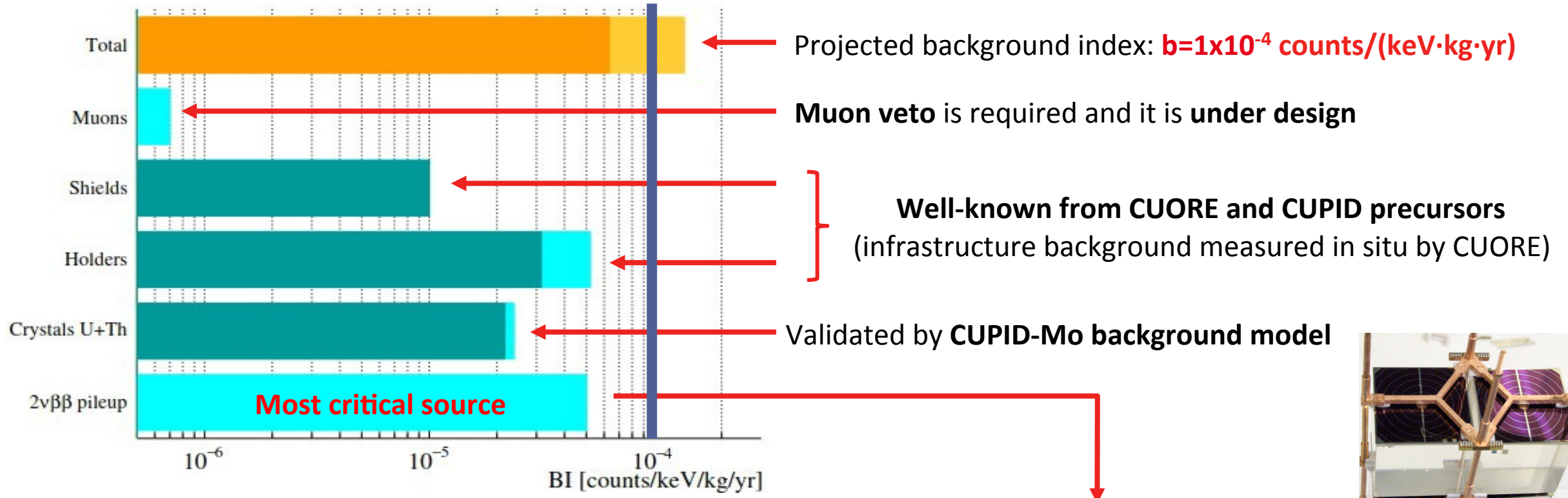
Light detector calibration

Irradiating the detector with a ^{133}Ba source we achieved an outstanding energy resolution:

0.71 keV FWHM at 356 keV



CUPID background



Random coincidence of $2\nu 2\beta$ events
($T_{1/2}^{2\nu 2\beta} = 7.1 \times 10^{18}$ y $\rightarrow$ 3 mHz in a CUPID crystal)
Pulse shape discrimination to reject pile up
Exploit faster light channels

- CUPID light-detector baseline: standard light detector with signal enhanced x10 by **Neganov-Trofimov-Luke effect** (IJCLab, Orsay)
Required performance demonstrated on 8 light detectors in Canfranc
Baseline energy resolution ~ 6 eV RMS – Risetime ~ 0.5 ms
 $\rightarrow$ Correspond to pile-up resolution of ~ 0.17 ms (CUPID goal)
- Alternative option for light detectors: TES readout (Berkeley)

CUPID sensitivity

Energy resolution: **5 keV FWHM**

Background index: **1×10^{-4} counts/(keV·kg·y)**

Livetime: **10 y**

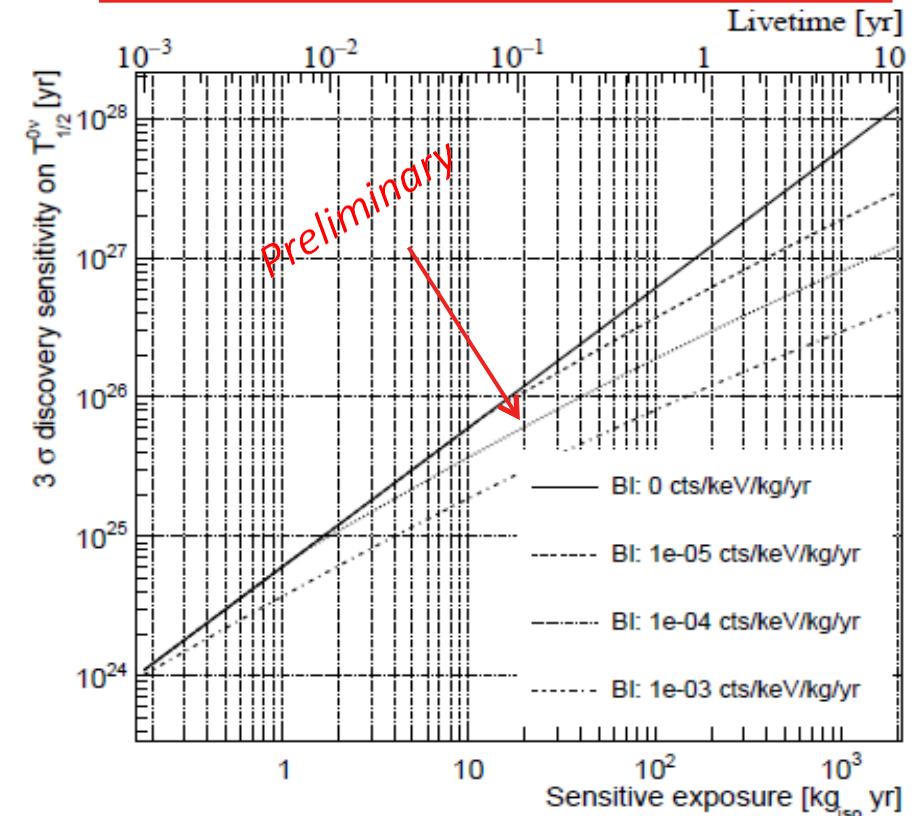
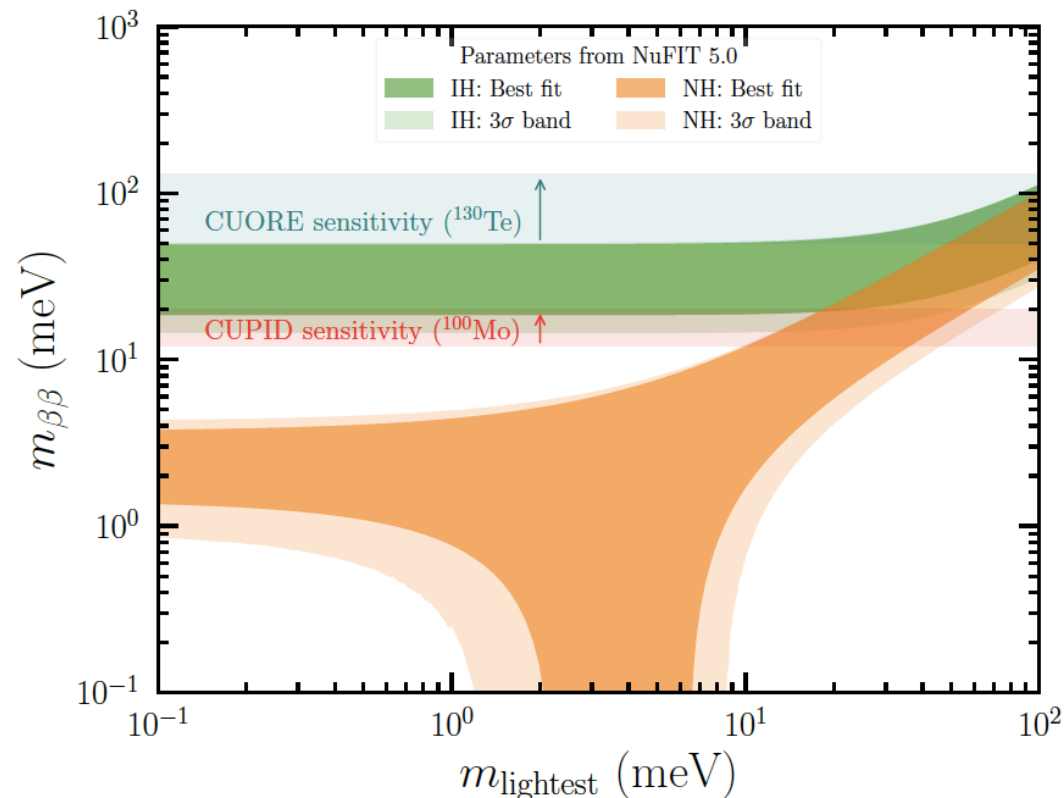


Half-life exclusion sensitivity

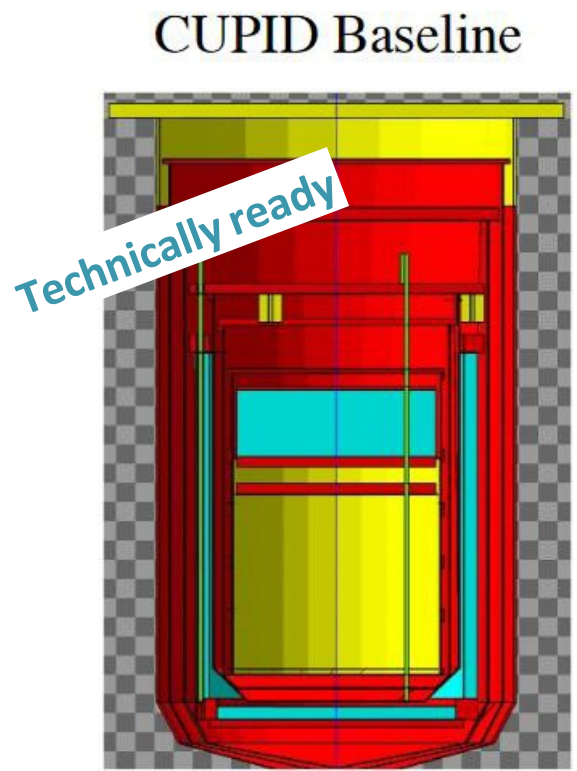
$1.4 \times 10^{27} \text{ y} - m_{\beta\beta} < 10\text{-}17 \text{ meV}$

Half-life 3σ discovery sensitivity

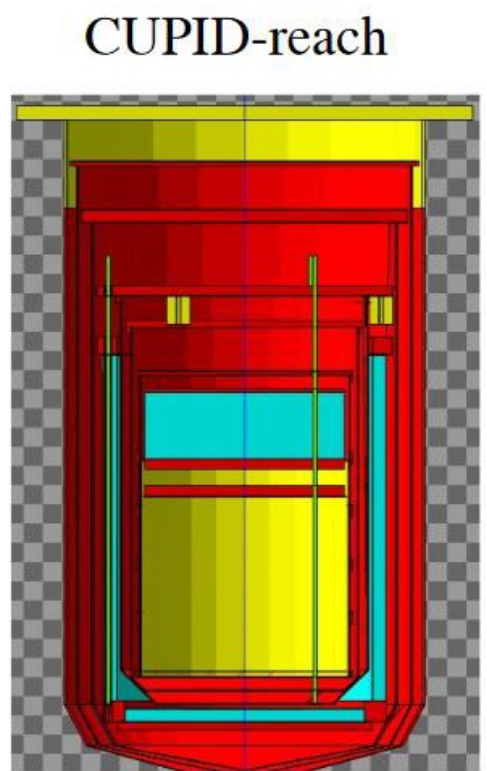
$1 \times 10^{27} \text{ y} - m_{\beta\beta} < 12\text{-}20 \text{ meV}$



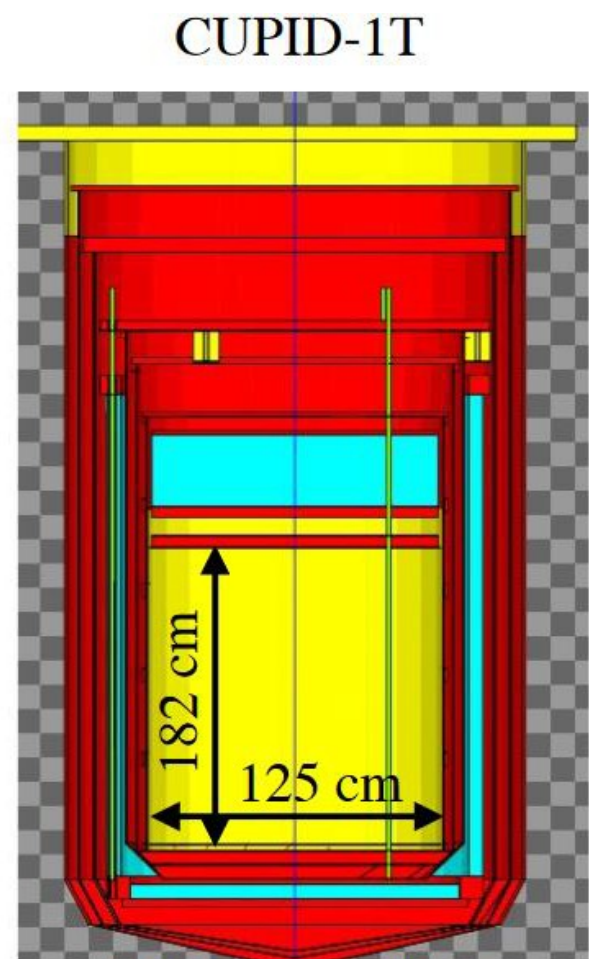
CUPID sensitivity: a phased approach



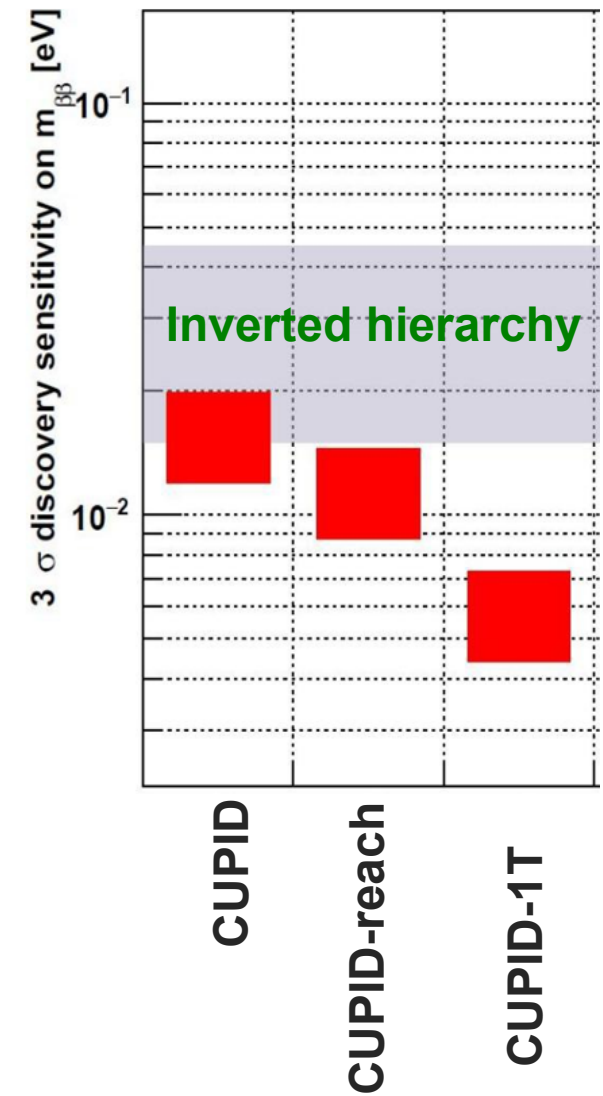
240 kg of ^{100}Mo
CUORE cryostat
Bkg 1×10^{-4} ckky
Excl. sensitivity:
 $T_{1/2} > 1.4 \times 10^{27} \text{ y}$



240 kg of ^{100}Mo
CUORE cryostat
Bkg 2×10^{-5} ckky
Excl. sensitivity:
 $T_{1/2} > 2.2 \times 10^{27} \text{ y}$

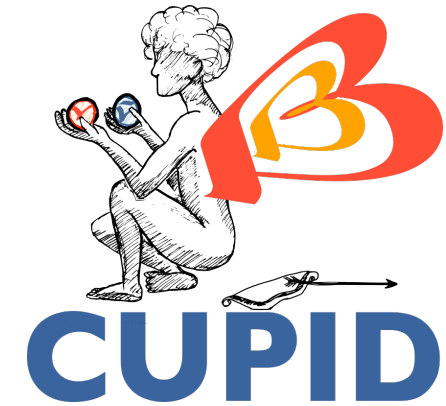


1000 kg of ^{100}Mo
New cryostat
Bkg 5×10^{-6} ckky
Excl. sensitivity:
 $T_{1/2} > 9.1 \times 10^{27} \text{ y}$



Conclusions

- The **CUPID infrastructure** already **exists** (**CUORE** cryostat, **LNGS**, Italy)
- **Scintillating bolometer Li_2MoO_4 technology** demonstrated in **CUPID-Mo**
- **Crystallization** and **enrichment** at large scale are **possible and demonstrated**
- The **performance** of the basic tower and of the single module are under test with promising results
- **Data-driven background model** indicates **$b \sim 10^{-4} \text{ counts}/(\text{keV} \cdot \text{kg} \cdot \text{y})$**
- **Fully explore the inverted ordering region**
→ down to $m_{\beta\beta} = 10 \text{ meV}$ for the most favorable nuclear model
- **The collaboration is working on getting ready for CUPID**
- On a longer time scale, **detector mass scaling** and potential **multi-isotope approach**



Related CUORE/CUPID/CUPID-Mo/CUPID-0



Recent progress on BSM and dark matter searches in CUORE

Alberto Ressa

Aug 28, 2023, 6:30 PM (#395)

First results from the CUORE background model

Stefano Ghislandi

Aug 29, 2023, 4:45 PM (Audimax)

Denoising Algorithms for the CUORE Experiment

Kenny Vetter

Aug 30, 2023, 3:30 PM (#216)

Impact of marine macroseisms on the response of the CUORE cryogenic calorimeters

Simone Quitadamo

Aug 30, 2023, 3:30 PM (#233)

Analysis techniques for the search of neutrinoless double-beta decay of Te-130 with CUORE

Krystal Alfonso

Aug 30, 2023, 3:30 PM (#557)

Final results of the CUPID-Mo $0\nu\beta\beta$ experiment

Léonard Imbert

Aug 29, 2023, 2:00 PM (Audimax)

Latest results from the CUORE experiment

Krystal Alfonso

Aug 29, 2023, 2:15 PM (Audimax)

Backgrounds and sensitivity of the CUPID experiment

Pia Loaiza

Aug 30, 2023, 3:30 PM (#183)

Development of enhanced light detectors for CUPID experiment

Vladyslav Berest, Anastasiia Zolotarova

Aug 30, 2023, 3:30 PM (#372)

Novel techniques for thermal detectors and applications for rare events physics

Irene Nutini

Aug 30, 2023, 3:30 PM (#396)

The Progress of Superconducting Transition Edge Sensor (TES) for Photothermal Detection System in Cupid-China $0\nu\beta\beta$ Experiment

Shasha Lv et al.

Aug 30, 2023, 3:30 PM (#628)

BINGO: investigation of the Majorana nature of neutrinos at a few meV level of the neutrino mass scale

Vladyslav Berest

Aug 30, 2023, 5:45 PM (Audimax)

Final results of the CUPID-0 combined background model

Emanuela Celi and Lorenzo Pagnanini

Aug 30, 2023, 6:15 PM (Audimax)